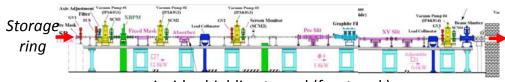
Cheiron school 2012, 25th Sep. 2012, SPring-8 X-ray beamline design II

Optics Engineering for x-ray beamline design

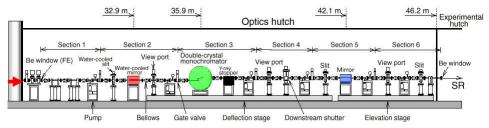
Haruhiko Ohashi JASRI / SPring-8

Light source(IDs/BM) X-ray beamline Radiation shield hutch ight source **Human safety** Interlock Brake Engine system Air-baas Machine protection Radiator Bodv Monochromator Tailoring x-rays (Power control Transmission Mirrors Gear User Interface A vehicle Steerina wheel **Application** Dashboard End station 3 Gear lever

Introduction "X-ray beamline looks complicated?"



inside shielding tunnel (front end)

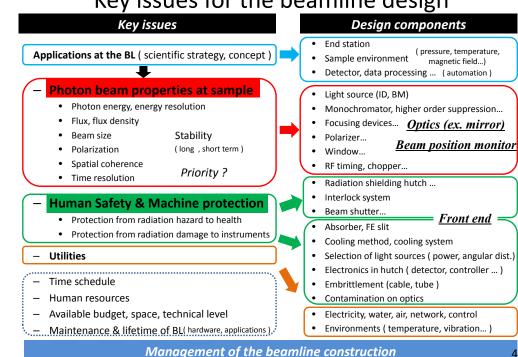


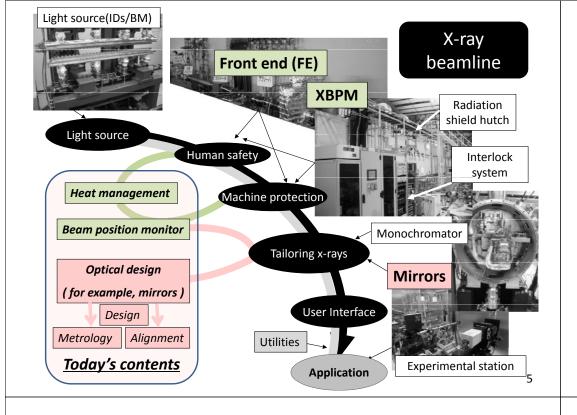
outside shielding tunnel (optics hutch)

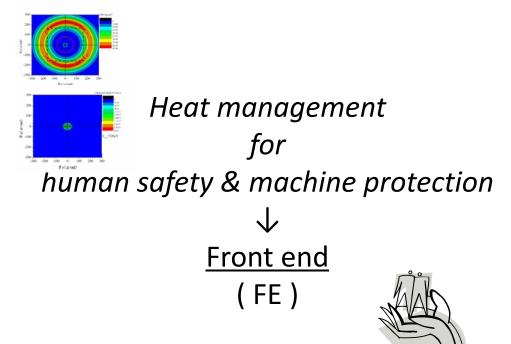
What function of each component?

2

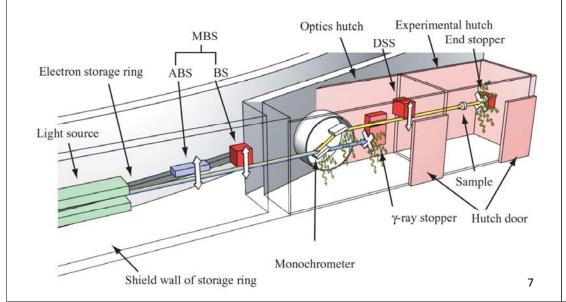
Key issues for the beamline design





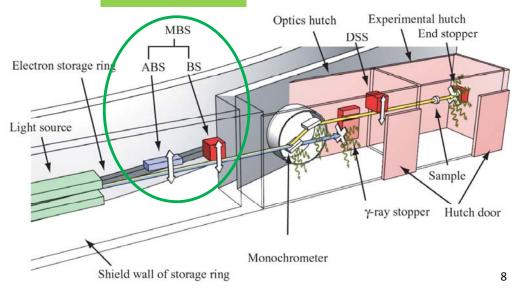


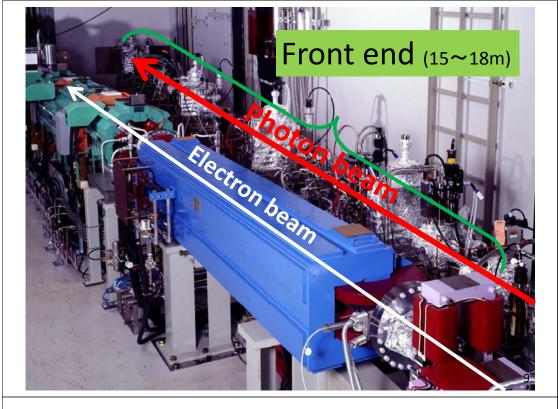
Beamline components for safety



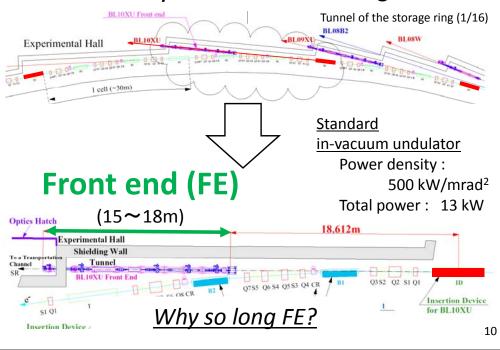
Beamline components for safety

Front end





Schematic Layout inside the SPring-8 Tunnel



Key functions & components of FE

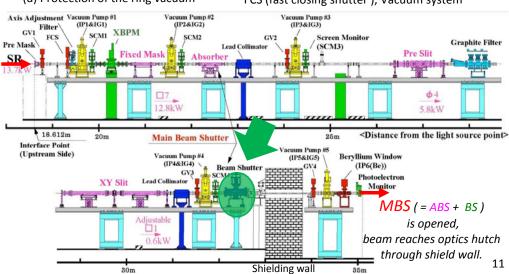
- (a) Shielding for human safety
- (b') Handling high heat load for optics
- (c) Monitoring the x-ray beam position
- (d) Protection of the ring vacuum

Beam shutter (BS), collimator (radiation shield)

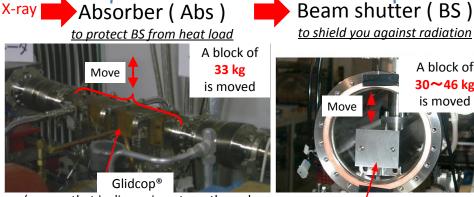
(b) Handling high heat load *for safety* Absorber, masks (to prevent BS from *melting*) XY slit, filters (to prevent optics from distorting)

XBPM (x-ray BPM), SCM (screen monitor)

FCS (fast closing shutter), Vacuum system



When we operate a main beam shutter (MBS), what happens?



(copper that is dispersion-strengthened with ultra-fine particles of aluminum oxide)

Heavy metal (alloy of tungsten) the thermal conductivity not so high

A block of

30~46 kg

is moved



After Abs is fully closed, BS is closed. After BS is fully opened, Abs is opened. The sequences are essential to keeping safety.



ABS and BS work on ways together to protect us from radiation when we enter the hutch. 12

Other key function is to handle high heat load for optics

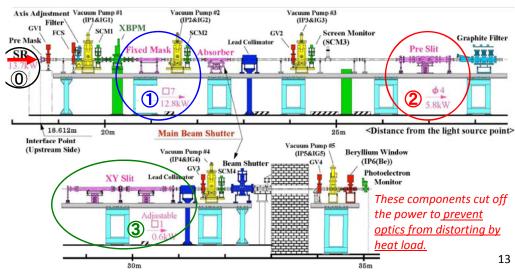
- (a) Shielding for human safety
- (b) Handling high heat load for safety
- (b') Handling high heat load for optics
- (c) Monitoring the x-ray beam position
- (d) Protection of the ring vacuum

Beam shutter (BS), collimator (radiation shield)
Absorber, masks (to prevent BS from melting)

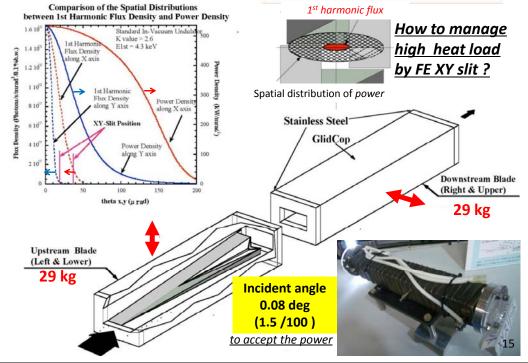
XY slit, filters (to prevent optics from distorting)

XBPM (x-ray BPM), SCM (screen monitor)

FCS (fast closing shutter), Vacuum system

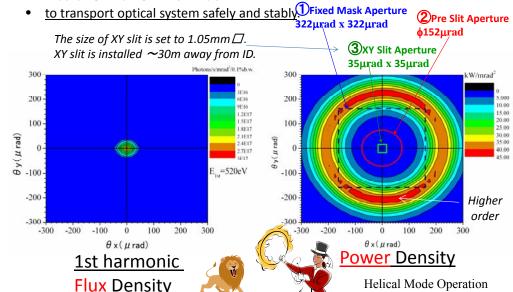


Slit: "Too much is as bad as too little"



FE: "For users to take lion's share"

- · Adding a spatial limitation to photon beam,
- supplying only a good quality part around the central axis of ID

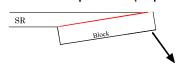


Handling Technology of high heat load

SPring-8 Standard In-Vacuum Undulator: 13.7kW, 550kW/mrad² at SPring-8

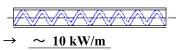
1. Grazed Angle Technology (Mask, Absorber, XY slit)

(1) Inclining absorbing surface to X-ray beam => Decrease of power density of per unit area



- (2) Applying the advanced material => Glid Cop
- (3) Enhancing the heat transfer coefficient of the cooling channel

=> Copper wire coil (SPring-8) Copper wire mesh (APS)



To increase the cooling ability within a more compact space

2. Volumetric Heating Technology
(Pre slit)

SR

Cooling Holder
(Copper)

Dissipating high surface heat flux in depth by utilizing a low-Z material, such as graphite or beryllium.

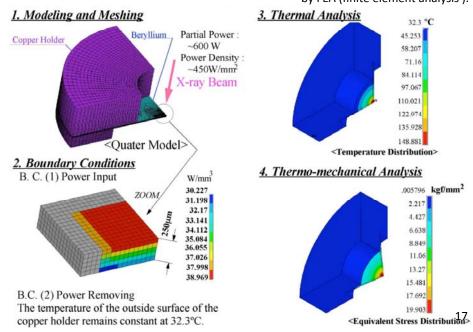
Developing the Volumetric Heating Mask

Target $\rightarrow \sim 5 \text{ kW/0.2m}$

at BL15XU 14

Simulation: "better safe than sorry"

For instance, the distributions of temperature and stress of Be window at FE can be calculated by FEA (finite element analysis).



Key issues of FE design

1. There exists a category of the beamline front ends.

They have their proper functions, proper missions based on the principles of human radiation safety, vacuum protection, heat-load and radiation damage protection of themselves.

They have to deal with every mode of ring operation and every mode of beamline activities.

- 2. Any troubles in one beamline should not make any negative effect to the other beamlines.
- 3. Strongly required to be a reliable and stable system.

We have to adopt key technologies which are reliable, stable and fully established as far as possible.

Higher the initial cost, the lower the running cost from the long-range cost-conscious point of view.

18

Monitoring stability of photon source

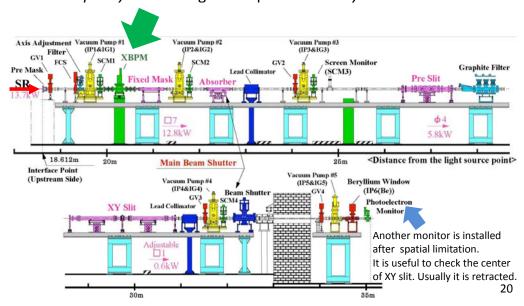


X-ray beam position monitor (XBPM)



Where is XBPM installed?

XBPM is installed before any spatial limitation. You hardly find it. It is *quietly* monitoring beam position *at any time*.

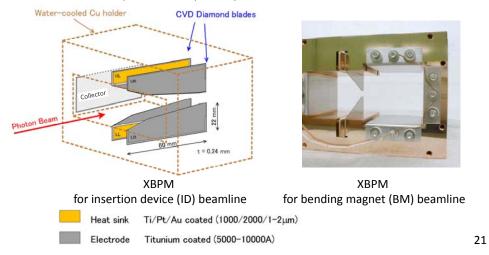


Revised Structure of XBPM's detector head

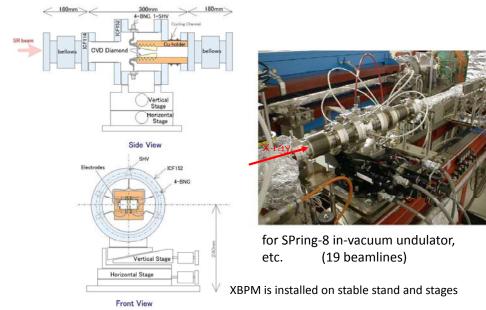
(Photo-emission type)

- Four blades are placed in parallel to the beam axis to reduce heat load.
- CVD diamond is used because of excellent heat property

Electrons from each blade of Ti/Pt/Au on diamond emitted by <u>outer side of photon beam</u> The horizontal or vertical positions computed by each current



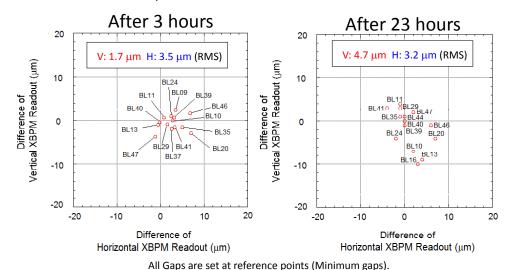
Fixed-blade style XBPM



22

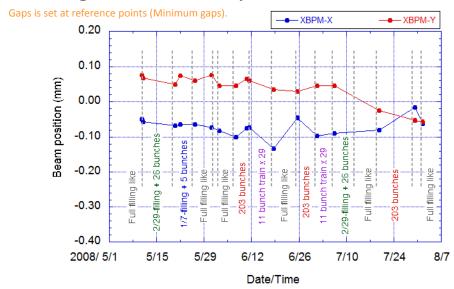
High stability of XBPM

As the stability is compared with other monitors outside wall, the stability of XBPM for 3 hours and 23 hours are measured.

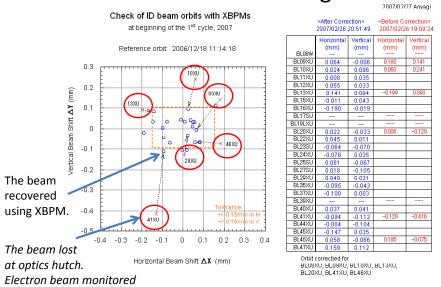


Stability of the XBPM is a few microns for a day under the same conditions (ID-gap, filling patter & ring current). 23

Long term stability of XBPM at BL47XU

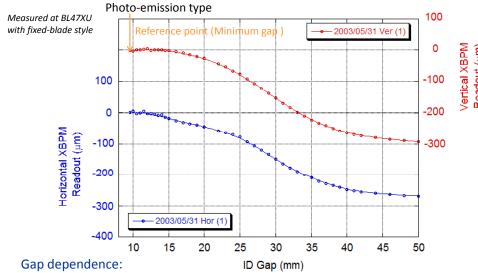


Orbit correction using XBPM



A fixed point observation of XBPM is helpful for a regular axis from ID.

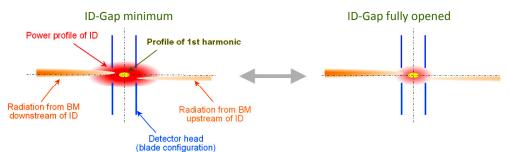
XBPM of ID-Gap dependence



 $\sim 100 \mu \text{m}$ for Gap : 9.6 $\sim 25 \text{ mm}$, $\sim 300 \mu \text{m}$ for Gap : 9.6 $\sim 50 \text{ mm}$

The position of the beam at optics hutch is fixed for changing ID gap.
What does the XBPM tell us?

What does the XBPM tell us?



Origin of ID-gap dependence of XBPM:

-XBPM of photo-emission type has energy dependence.

Radiation from ID changes drastically, but not from BMs (backgrounds)

- Backgrounds are asymmetric and usually offset.

1st harmonic: $6 \sim 18 \text{ keV}$,

Background: < several keV near beam axis of ID

XBPM depends on ID-gap, filling pattern & ring current.

The results of XBPM can be compared with the same condition.

Key issues of XBPM design

for high power undulator radiation in SPring-8

1. Dependence of ID gap, ring current, filling pattern

XBPM (photo-emission type) depends on these parameters.

2. High stability

25

XBPM has stability of microns for a day.

3. Resolution of x-ray beam position

- The resolution of micron order can be monitored.

Beam divergences are ~ 20 / 5 μrad (hor. / ver.), which correspond to beam sizes of ~ 400 / 100 μm (hor. / ver.) at XBPM position (20 m from ID).

4. Withstand high heat Load

- Blade of diamond

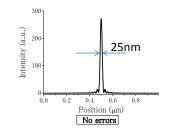
Max. power density is ~ 500 kW/mrad². Metal will melt immediately.

5. Fast Response

- Response time of < 1 msec needs for high frequency diagnostic.
- Simultaneous diagnostic over beamlines is important.

Ref. of XBPM: for example, H. Aoyagi et al., "High-speed and simultaneous photon beam diagnostic system using optical cables at SPring-8", AIP Conf.Proc.705-593 (2004).





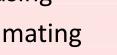
Tailoring x-rays to application



X-ray mirrors design, errors, metrology & alignment

The functions of x-ray mirrors

- Deflecting
- Low pass filter
- Focusing
- Collimating



- Separation from y-ray
- Branch / switch beamline
- Higher order suppression
- Micro- / nano- probe
- Imaging
- Energy resolution w. multilayer or crystal mono.



30

Design parameters of x-ray mirror

Requirement

the beam properties both of incident and reflected x-rays

(size, angular divergence / convergence, direction, energy region, power...)

We have to know well what kinds beam irradiate on the mirror.

Design parameters

Coating material : Rh, Pt, Ni ... (w/o binder , Cr), thickness

: multilayers (ML), laterally graded ML

Incident angle : grazing angle (mrad) How to select

Surface shape : flat, sphere, cylinder, elliptic ...

: adaptive (mechanically bent, bimorph)

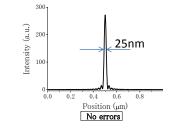
Substrate shape : rectangular, trapezoidal... Substrate size : length, thickness, width

: indirect or direct, water or LN₂... w/o cooling

Substrate material: Si, SiO2, SiC, Glidcop...

In addition,

some errors such as figure error, roughness...



Tailoring x-rays to application



X-ray mirrors

design, errors, metrology & alignment



How to select coating material and incident angle?

Reflectivity for grazing incident mirrors

$$R(\lambda, \theta, n) = \left| \frac{k_1 - k_2}{k_1 + k_2} \right|^2$$

$$k_1 = \frac{2\pi}{\lambda} \cos \theta, k_2 = \frac{2\pi}{\lambda} \sqrt{n^2 - \cos^2 \theta}$$

The complex index of refraction

Coating material (1) "the complex index of refraction"

The complex atomic scattering factor for the forward scattering

$$f = f_1 + if_2$$

The complex index of refraction

$$n = 1 - \delta - i\beta$$
 $E \propto e^{-i(\varpi t - kr)}$

$$E \propto e^{-i(\varpi t - kr)}$$

$$\begin{cases} \delta = \frac{Nr_0\lambda^2}{2\pi} f_1(\lambda) \\ \beta = \frac{Nr_0\lambda^2}{2\pi} f_2(\lambda) \end{cases}$$

$$\beta = \frac{Nr_0\lambda^2}{2\pi} f_2(\lambda)$$

	δ (×10 ⁻⁵)	β(×10 ⁻⁷)
Si	0.488	0.744
Quartz	0.555	2.33
Pt	3.26	20.7
Au	2.96	19.5

$$r_0 = \frac{e^2}{4\pi mc^2} = 2.82 \times 10^{-15} m$$
 $\beta = \frac{\mu\lambda}{4}$

$$\beta = \frac{\mu\lambda}{4\pi}$$

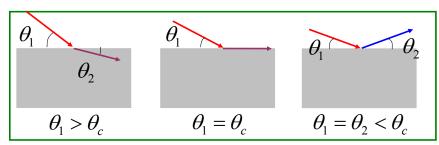
μ: linear absorption coefficient

N: Number of atoms per volume

Coating material (2)

"total reflection"

$$n_1/n_2 = \cos(\theta_1)/\cos(\theta_2)$$
 Snell's law



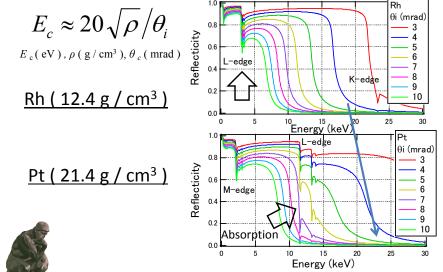
$$\cos(\theta_c) = n = 1 - \delta, \cos(\theta_c) \rightarrow 1 - \theta_c^2/2$$

$$\theta_c \cong \sqrt{2\delta} = 1.6 \times 10^{-2} \lambda \sqrt{\rho} = 20 \sqrt{\rho} / E$$

For example, θ_c (rad), ρ (g/cm³), λ (nm), E (eV) Rh ($\rho = 12.4 \text{ g/cm}^3$) $\lambda = 0.1 \text{nm}, \theta_c = 5.68 \text{ mrad}$

Coating material (3): "cut off, absorption"

The cut off energy of total reflection *Ec*



Cut off energy, absorption \rightarrow incident angle

Opening of the mirror, length, width of mirror, power density 36

35

Atomic scattering factors, Reflectivity

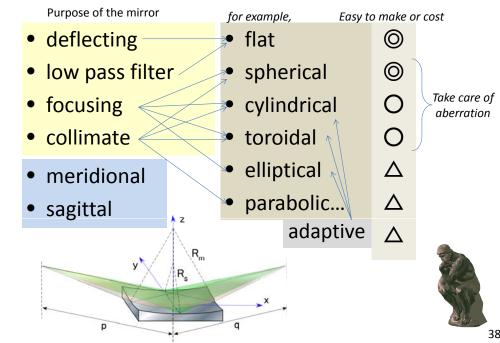
You can easily find optical property in "X-Ray Data Booklet" by Center for X-ray Optics and Advanced Light Source,
Lawrence Berkeley National Lab.
In the site the reflectivity of x-ray mirrors can be calculated.

http://xdb.lbl.gov/



Many thanks to the authors!

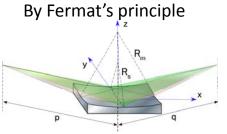
Surface shape (1)



Surface shape (2) radius and depth

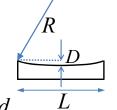
$$R_{m} = \frac{2}{(1/p + 1/q)\sin(\theta_{i})}$$

$$R_{s} = \frac{2\sin(\theta_{i})}{(1/p + 1/q)} = R_{m}\sin^{2}(\theta_{i})$$



For parallel beam $q \rightarrow \infty, 1/q = 0$

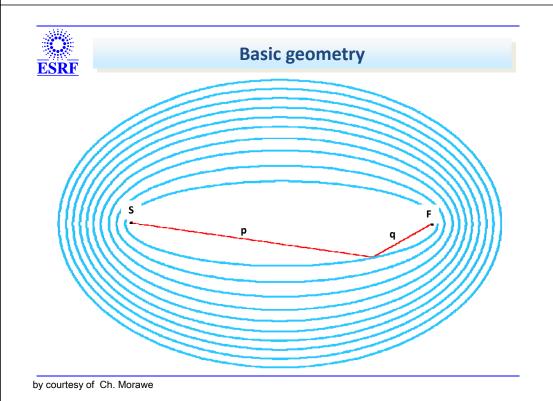
Depth at the center $_{D=R}-\sqrt{R^{2}-\left(\frac{L}{2R}\right)^{2}}\approx\frac{L^{2}}{8R}$ For example,

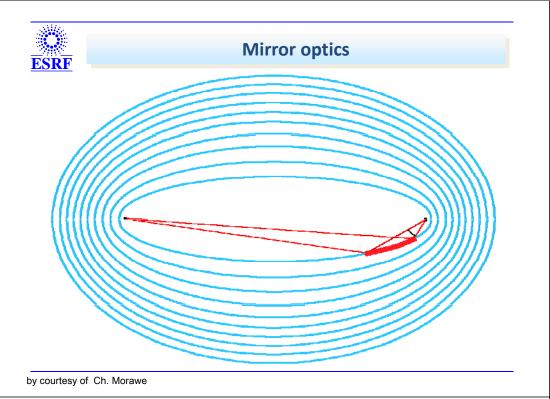


 $p=15 \sim 50m$, $q=5 \sim 20m \theta_i=1 \sim 10mrad$

 $R_m = 0.1 \sim 10 \text{ km}, R_s = 30 \sim 100 \text{ mm}$

 $R=1 \text{ km}, L=1\text{m} \rightarrow D=125 \text{ }\mu\text{m}$



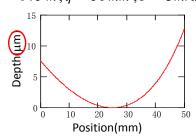


Surface shape (3) elliptical

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

For example,

$$p = 975 \, m, q = 50 \, mm, \theta = 3 \, mrad$$



 $a = \frac{p+q}{2}$ $b = \frac{1}{2}\sqrt{2pq(1-\cos(2\theta))}$

$$x_0 = \frac{p^2 - q^2}{2\sqrt{p^2 + 2pq\cos(2\theta) + q^2}}$$

$$y_0 = -b\sqrt{1 - \frac{{x_0}^2}{a^2}}$$

$$u = \frac{\frac{\sigma}{a} \times x_0}{\sqrt{a^2 - x_0^2}}$$

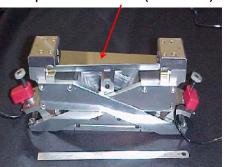
 $z(s) = -\cos(u) \times b\sqrt{1 - \left(\frac{s \times \cos(u) + x_0}{a}\right)^2} + s \times \cos(u)\sin(-u)$

Precise fabrication is difficult.

(Ref *)

Elliptical mirror mechanically bent using trapezoidal substrate

Trapezoidal mirror (L170mm)



Dynamically bent KB mirror at ESRF

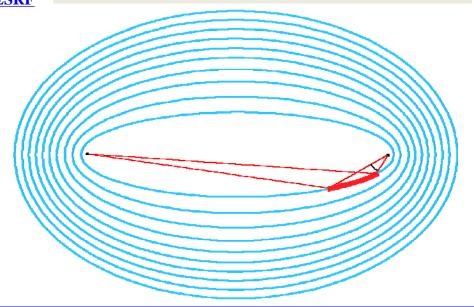
Trapezoidal mirror (L540mm)



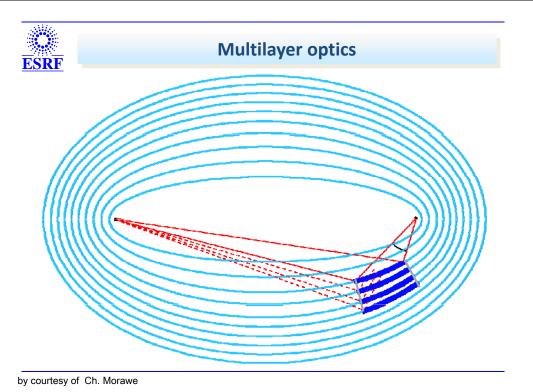
Long bent focusing mirror at SPring-8

ESRF

Mirror optics



^{*} M.R Howells et al., "Theory and practice of elliptically bent X-ray mirrors", Optical Eng. **39**, 2748 (2000).





X-ray multilayer characterization

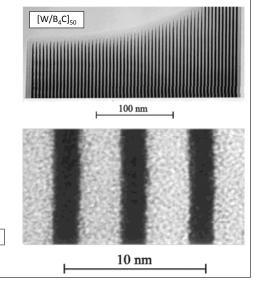
Transmission electron microscopy (TEM)

- Fabrication errors
- Roughness evolution
- Crystallinity
- Interface diffusion



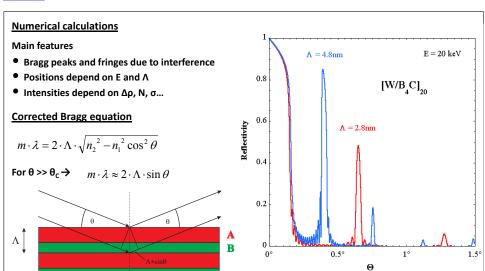
Complementary to x-ray measurements !

R. Scholz, MPI Halle, Germany



ESRF

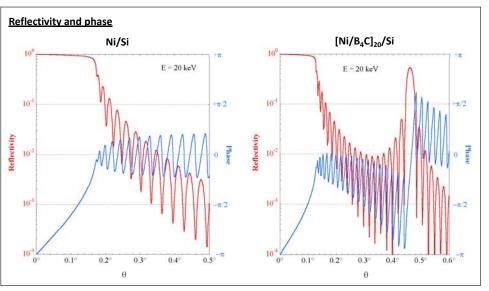
X-ray multilayer reflectivity



by courtesy of Ch. Morawe



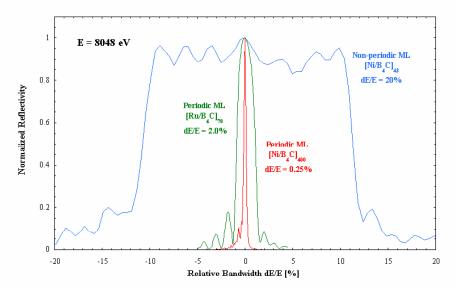
X-ray reflectivity



by courtesy of Ch. Morawe



Energy resolution of multilayers



by courtesy of Ch. Morawe

Design parameters of x-ray mirror

Requirement

the beam properties both *of incident and reflected x-rays* (size, angular divergence / convergence, direction, energy region, power...)

Design parameters

• Coating material : Rh, Pt, Ni ... (w/o binder , Cr), thickness

: multilayers (ML), laterally graded ML

• Incident angle : grazing angle (mrad)

• Surface shape : flat, sphere, cylinder, elliptic ...

: adaptive (mechanically bent, bimorph)

Substrate shape : rectangular, trapezoidal...Substrate size : length, thickness, width

• w/o cooling : indirect or direct, water or LN₂...

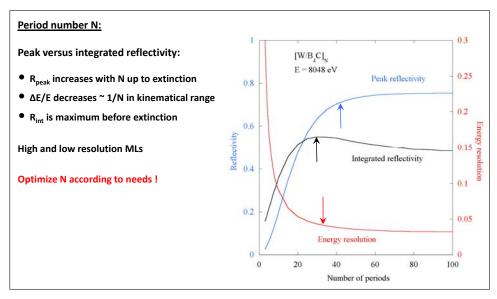
• Substrate material: Si, SiO2, SiC, Glidcop...

In addition,

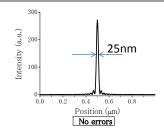
some errors such as figure error, roughness...

ESRF

X-ray multilayer design



by courtesy of Ch. Morawe



Tailoring x-rays to application



X-ray mirrors

design, errors, metrology

& alignment



"An actual mirror has some errors."

The tolerance should be specified to order the mirror

- Roughness
- Density of coating material
- Radius error
- Figure error

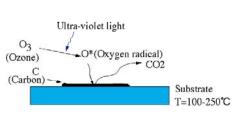
- Reflectivity
- Beam size
- Distortion
- Deformation ...

The cost (price and lead time) depends entirely on tolerance. We must consider or discuss how to measure it.

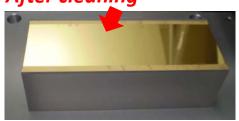
- Deformation by self-weight, coating and support ...
- Figure error of adaptive mechanism
- Misalignment of mirror
- Stability of mirror's position (angle)
- Deposition of contamination by use
- Decomposition of substrate by use
- Environment
- Stages
- Cooling system ...

Contamination and removal

before



After cleaning



Advantage of UV/ozone cleaning

- 1. Low Damage
- 2. Contamination-free
- 3. Non-contact

UV / ozone cleaning

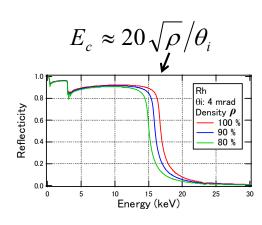
It takes from 10 min to a few hours.

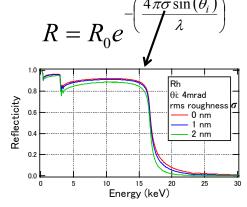
54

56

Errors (1)

"Density $\boldsymbol{\rho}$ and surface roughness $\boldsymbol{\sigma}$ "

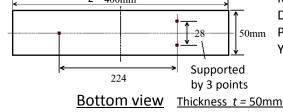




Coating on sample wafer at the same time is helpful to evaluate the density and roughness.

Errors (2)

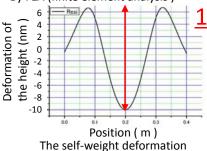
"the self-weight deformation"



Material SiO₂ $2.2 \, g / cm^3$ Density Poisson's ratio 0.22 Young's modulus E = 70 Gpa

 $D \propto \frac{L^4}{F \times t^3}$

By FEA (finite element analysis)



on the surface of mirror (SiO₂)

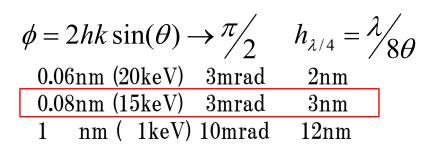
16.7 nm PV

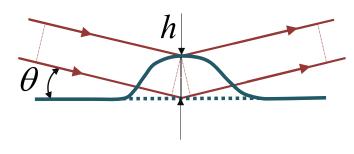
This value is larger than figure error by Rayleigh's rule. (→See next page)

Improvement for nano-focusing

- a) Substrate \rightarrow Si (E \sim 190 GPa)
- b) Optimization of supporting points and method
- c) Figuring in consideration of the deformation



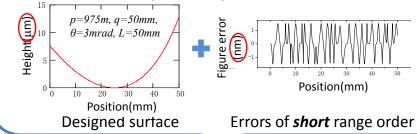




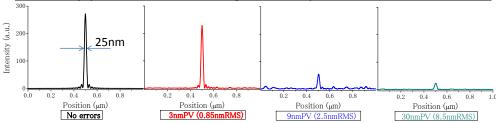
57

Errors (3b)

" estimation by wavefront simulation"



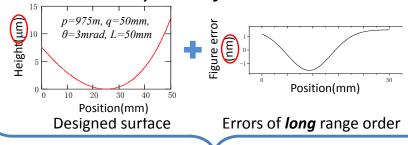
Intensity profiles of focusing beam by wavefront simulation



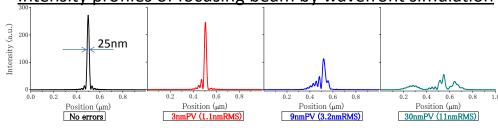
Errors of short range order decreases intensity. → Roughness

Errors (3c)

" estimation by wavefront simulation"

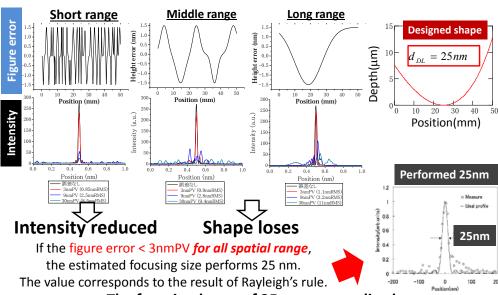


Intensity profiles of focusing beam by wavefront simulation



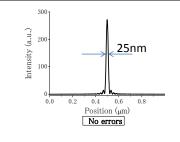
Errors of long range order loses shape.

" estimation by wavefront simulation"



The focusing beam of 25 nm was realized using high precision mirror with figure error of 3 nm PV

*H. Mimura, H. Yumoto, K. Yamauchi et.al, Appl. Phys. Lett. **90**, 051903 (2007).



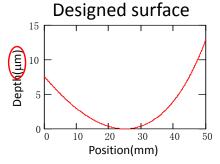
Tailoring x-rays to application

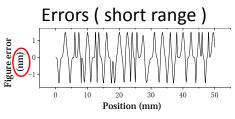


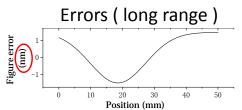
X-ray mirrors

design, errors, metrology & alignment

How to evaluate the errors?







62

Metrology instruments for x-ray optics

Long /

middle

~ 0.1 m,

0.1 mm

Field of view, lateral resolution

Short / middle Short ~10 mm. ~10 µm, 1 µm 0.1 nm

Roughness

Scanning probe

microscope

z (0.1nm)

Roughness, figure

Scanning white light interferometer z (0.1nm)

Fizeau interferometer

z (0.1nm)

Long Trace Profiler (LTP)

middle

Slope

Long /

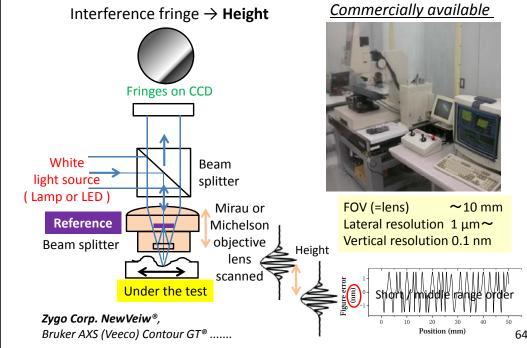
~1m,

1 mm

Vertical resolution (rms)

slope (0.1urad)

Scanning white light interferometer

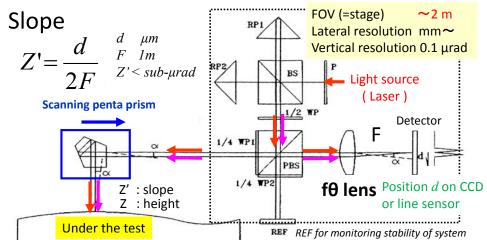


Fizeau interferometer Interference pattern → **Height** Commercially available Monochromatic Zygo Corp. VeriFire®, point light source 4DS technologies, FujiFILM Beam splitter **Fizeau fringes on CCD** FOV (=reference) ~0.1 m Lateral resolution ∼0.1 mm Vertical resolution 0.1 nm Collimator Reference Long \middle range order Cavity Under the test Not easy to measure large mirro 65

Long trace profiler (LTP)

Direction of laser reflected on the surface \rightarrow **Slope**

<u>Homemade</u>

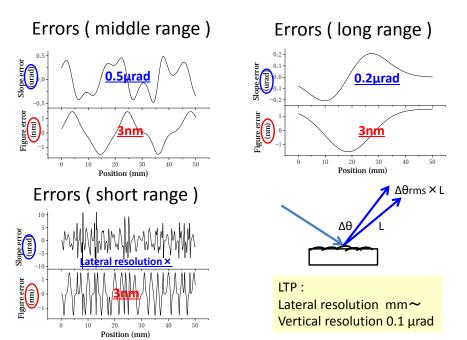


Easy to measure slope of sub-μrad on large mirror by NO reference Many kinds of LTPs are developing among SR facilities.

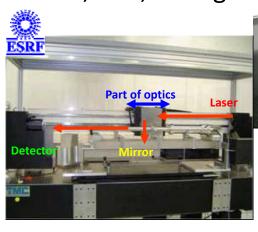
For example, S. Qian, G. Sostero and P. Z. Takacs, Opt. Eng. 39, 304-310 (2000).

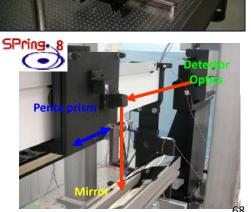
66

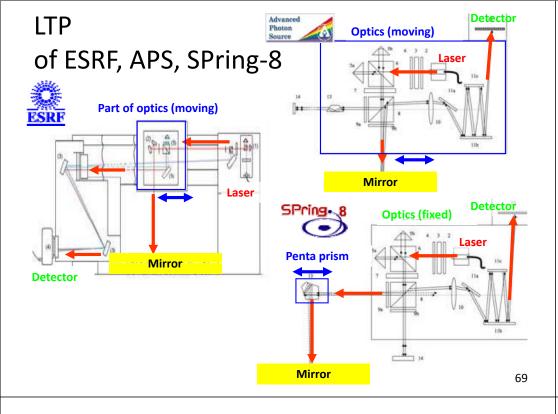
Figure error and slope error



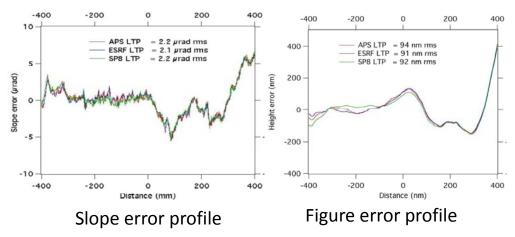
LTP of ESRF, APS, SPring-8







Round robin measurement of 1m-long toroidal mirror

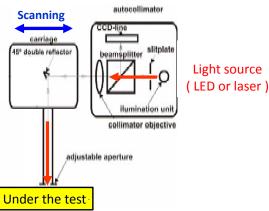


L. Assoufid, A. Rommeveaux, H. Ohashi, K. Yamauchi, H. Mimura, J. Qian, O. Hignette, T. Ishikawa, C. Morawe, A. T. Macrander and S. Goto, SPIE Proc. 5921-21, 2005, pp.129-140.

70

Nanometer Optical Component Measuring Machine (NOM) @HZB

Autocollimator → **Slope** Homemade



F. Siewert et al.: "The Nanometer Optic Component Measuring Machine: a new Sub-nm Topography "SRI 2003, AIP Conf. Proc.

Stitching interferometer for large mirror Homemade

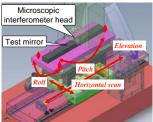
MSI

RADSI

(micro-stitching interferometer)

(relative angle determinable stitching interferometer)





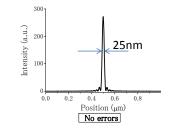




Collaboration with Osaka Univ., JTEC and SPring-8
H. Ohashi et al., Proc. Of SPIE **6704**, 670405-1 (2007)

Height error of wide range order for a long and aspherical mirror with $1\mu m$ of lateral and 0.1 nm of vertical resolution.

Necessity is the mother of invention.



Tailoring x-rays to application

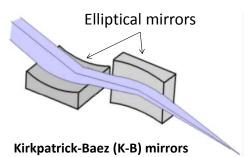


X-ray mirrors

design, errors, metrology & alignment



Introduction of KB mirrors



In 1948, P. Kirkpatrick and A. V. Baez proposed the focusing optical system.

P. Kirkpatrick and A. V. Baez, "Formation of Optical Images by X-Rays", J. Opt. Soc. Am. 38, 766 (1948).

Advantages

- •Large acceptable aperture and High efficiency
- No chromatic aberration
- Long working distance

Disadvantages

- Difficulty in mirror alignments
- Difficulty in mirror fabrications
- Large system

Suitable for x-ray nano-probe

74

Overview of x-ray focusing devices

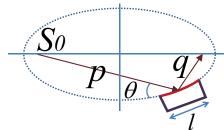
Diffraction	focus size, focal length [energy]	energy range	aberration -coma -chromatic -figure error
Fresnel Zone Plate	12 nm, f = 0.16 mm [0.7 keV], 30 nm, f = 8 cm [8 keV]	soft x-ray hard x-ray	-coma small -chromatic exist -figure error small
Sputter sliced FZP	0.3 µm, f = 22 cm [12.4 keV], 0.5 µm, f = 90 cm [100 keV]	8-100 keV	-coma small -chromatic exist -figure error large→small
Bragg FZP	2.4 µm, f = 70 cm [13.3 keV]	mainly hard x-ray	-coma small -chromatic exist -figure error small
Multilayer Laue Lens	16 nm(1D), f = 2.6 mm [19.5 keV], 25nm × 40nm, f=2.6mm,4.7mm [19.5 keV]	mainly hard x-ray	-coma large -chromatic exist -figure error small

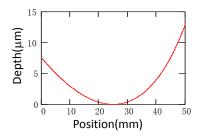
Refraction	focus size, focal length [energy]	energy range	aberration -coma -chromatic -figure error
Pressed Lens	1.5 µm, f = 80 cm [18.4 keV], 1.6 µm, f = 1.3 m [15 keV]	mainly hard x-ray	-coma small -chromatic exist -figure error large
Etching Lens	47nm × 55nm, f = 1cm, 2cm [21 keV]	mainly hard x-ray	-coma small -chromatic exist -figure error small

_			
Reflection			
Kirkpatrick-Baez Mirror	36nm × 48nm, f=15cm,25cm [15 keV], 7 nm(1D), f=7.5cm [20 keV]	soft x-ray hard x-ray	-coma large -chromatic not exist -figure error small
Wolter Mirror	0.7 μm, f = 35 cm [9 keV]	<10 keV	-coma small -chromatic not exist -figure error large
X-ray Waveguide	95 nm, [10 keV]	soft x-ray hard x-ray	-coma large -chromatic not exist -figure error large

How small is x-ray focused?

For example, by elliptical mirror





Geometrical size

<u>Diffraction limited size</u>(FWHM)

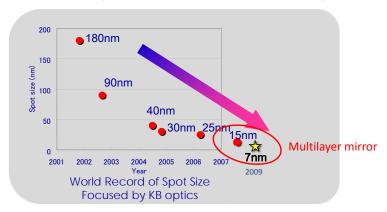
$$d_{DL} = \lambda \times \frac{0.88q}{1\sin(\theta)}$$

 $p=975\,m,q=50\,mm$, $\theta=3mrad$, $l=50\,mm$, $\lambda=0.083\,nm$, $S_0=100\,\mu m$

Mag. = 1 / 19500
$$d_G = 5nm < d_{DL} = 25nm$$

The opening of the mirror restricts the focused size even if magnification is large. 76

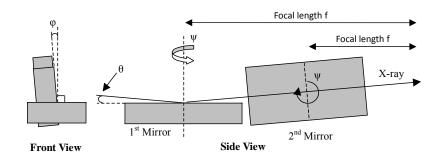
Nano-focusing by KB mirror History since the century



World Record of spot size is **7** nm (by Osaka Univ., in 2009 *). Routinely obtained spot size is up to **30** nm.

Ref*: H. Mimura et al., "Breaking the 10 nm barrier in hard-X-ray focusing", Nature Physics 6, 122 (2010)-7

Difficulty in mirror alignments



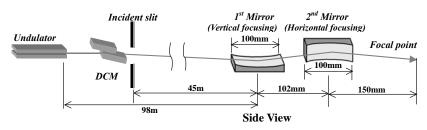
Positioning two mirrors is difficult because there are at least 7 degree of freedom.



It is difficult to use KB mirrors.

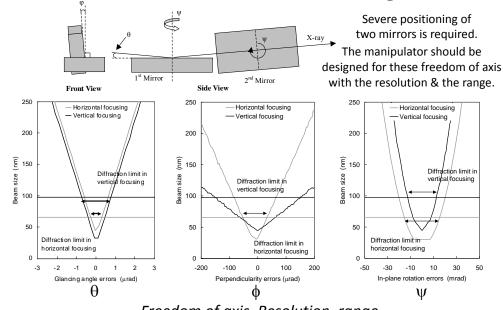
78

KB optics installed in BL29XU-L



	1st Mirror	2 nd Mirror
Glancing angle (mrad)	3.80	3.60
Mirror length (mm)	100	100
Mirror aperture (μm)	382	365
Focal length (mm)	252	150
Demagnification	189	318
Numerical aperture	0.75x10 ⁻³	1.20x10 ⁻³
Coefficient a of elliptic function (mm)	23.876 x 10 ³	23.825 x 10 ³
Coefficient b of elliptic function (mm)	13.147	9.609
Diffraction limited focal size (nm, FWHM)	48	29

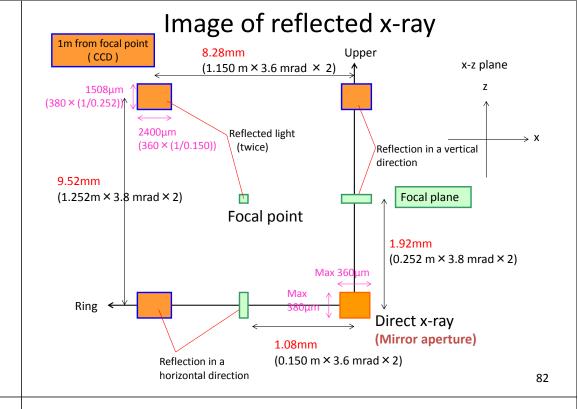
Tolerance limits of mirror alignments



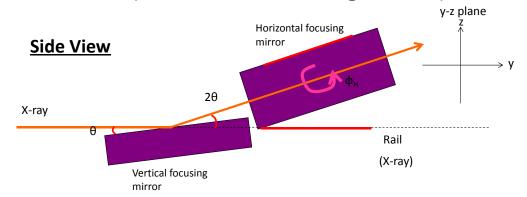
<u>Freedom of axis, Resolution, range</u>

Ref: S. Matsuyama, H. Mimura, H. Yumoto et al., "Development of mirror manipulator for hard-x-ray nanofocusing at sub-50-nm level", Rev. Sci. Instrum. **77**, 093107 (2006).

Image on X-ray CCD camera L L $L \times 2\theta = x$ $\theta = \frac{x}{2L}$



Alignment of in-plane rotation (Horizontal focusing mirror)



 θ : 3.8mrad \rightarrow 2 θ : 7.6mrad

Reflected angle of vertical-focusing mirror needs to be considered, in the alignment of in-plane rotation of horizontal-focusing mirror.

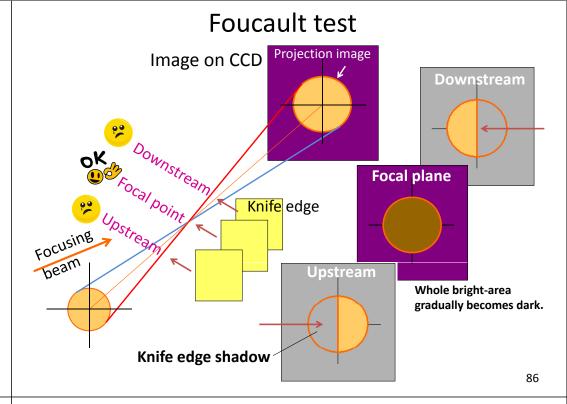
Alignment of incident angle

- Foucault test
 Rough assessment of focusing beam profile.

 This method is used for seeking focal point.
- Wire (Knife-edge) scan methodFinal assessment of focusing beam profile.

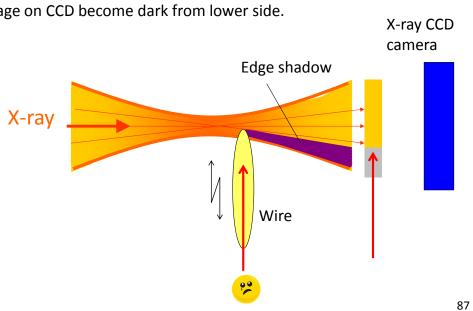
Precise adjustment of the glancing angle and focal distance is performed until the best focusing is achieved, while monitoring the intensity profile.

Alignment of incident angle X-ray X-ray X-ray X-ray X-ray X-ray X-ray CCD camera



Foucault test 1

Wire is at downstream of focal point.
Image on CCD become dark from lower side.



Foucault test 2

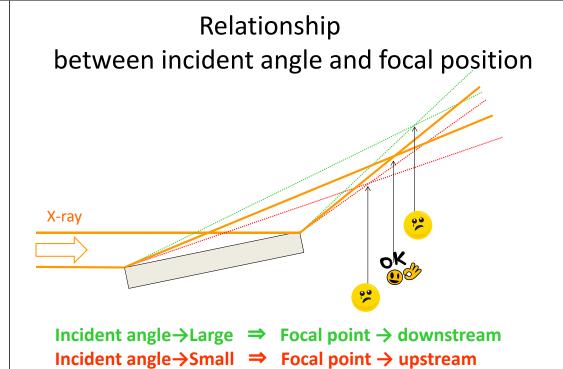
85

Wire is at upstream of focal point.
Image on CCD become dark from upper side.

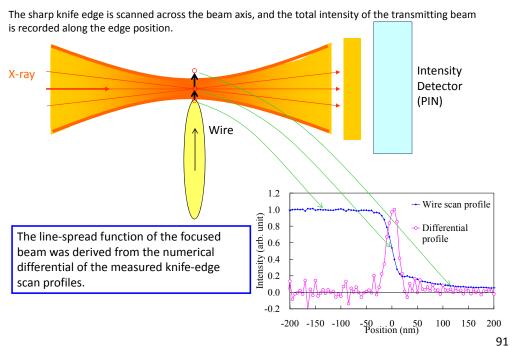
X-ray CCD camera

Wire

Wire is at the focal point. Whole bright-area gradually becomes dark. X-ray CCD camera X-ray Wire



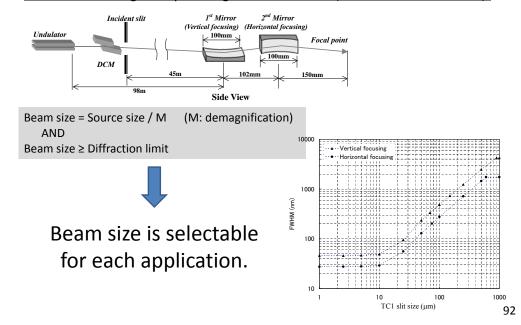
Wire (Knife-edge) scan method for measuring beam profiles



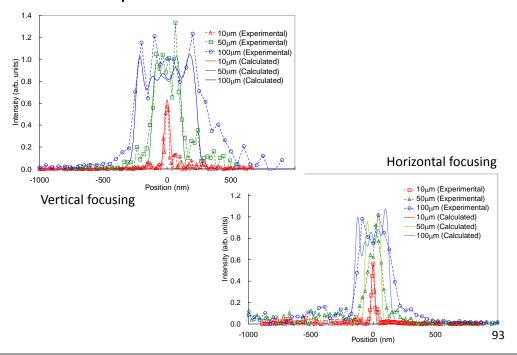
Relationship between Beam size and Source size

90

Beam size changes depending on source size (or virtual source size).



Relationship between Beam size and Source size

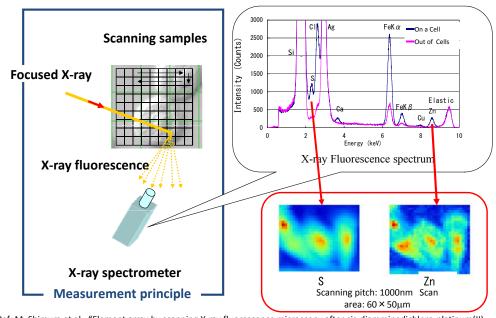


Key issues of x-ray mirror design

- To select the functions of x-ray mirror
 - Deflecting, low pass filtering, focusing and collimating → Shape of the mirror
- To specify the incident and reflected beam properties Energy range, flux
 - \rightarrow absorption, cut off energy \rightarrow coating material \rightarrow incident angle The beam size and the power of incident beam
 - → opening of the mirror, incident angle
 - \rightarrow absorbed power density on the mirror \rightarrow w/o cooling, substrate
 - Angular divergence / convergence, the reflected beam size
 - → incident angle, position of the mirror (source, image to mirror) Direction of the beam
 - → effect of polarization, self-weight deformation
- To specify the tolerance of designed parameters
 - Roughness, density of coating material, radius error, figure error The cost (price and lead time) depends entirely on the tolerance.
- *To consider the alignment*
 - The freedom, resolution and range of the manipulator



Scanning X-ray Fluorescence Microscope: SXFM



Ref: M. Shimura et al., "Element array by scanning X-ray fluorescence microscopy after cis-diamminedichloro-platinum(II) treatment", Cancer research 65, 4998 (2005).

Key issues for the beamline design

Key issues

Which application is the most important at the BL? Can you specify who uses the property at the BL?

- Photon energy, energy resolution
- Flux, flux density

Beam size

The higher, the better? The smaller, the better ? More is NOT always better

- Polarization
 - Spatial coherence Simplify the property.
- Time resolution Get your priorities right.
- Time schedule
- Human resources
- Available budget, space, technical level
- Maintenance for keeping performance
- Lifetime of the BL (hardware and application)

What to include or not? What to develop or not?

Design components

- · End station
- (pressure, temperature, • Sample environment
- · Detector, data processing ... (automation)
- Light source (ID, BM)
- · Monochromator, higher order suppression...
- · Focusing devices...
- Polarizer...

Stability enough

to measure

Safety first!

- Window...
- · RF timing, chopper...
- · Radiation shielding hutch ...
- Interlock system
- Beam shutter...
- Absorber, FE slit
- · Cooling method, cooling system
- Selection of light sources (power, angular dist.)
- · Electronics in hutch (detector, controller ...)
- · Embrittlement (cable, tube)
- Contamination on optics
- Electricity, water, air, network, control
- Environments (temperature, vibration...

Simple and clear design to accelerate your research

Ongoing x-ray beamline

X-ray beamline looks complicated, but the function of each component is simple.

To specify the beam properties is to design the beamline.

New x-ray beamline for next generation light source such as XFEL is newly constructed.

The components for heat management, x-ray beam monitors and x-ray optics including metrology are newly developed to perform the beam properties.

Challenges at XFEL beamline:

coherence preservation
wavefront disturbance or control
at wavelength technique
ultra-short & high intense pulse
high stability
shot-by-shot diagnosis of x-rays
timing control of x-ray pulse
synchronization with other source ...

Acknowledgment

- S. Takahashi (front end), H. Aoyagi (XBPM),
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- S. Matsuyama, H. Yumoto, H. Mimura, K. Yamauchi (ultimate focusing mirror, alignment),
- C. Morawe (multilayer) ESRF
- S. Goto and T. Ishikawa

97

Thank you for your kind attention.

Enjoy Cheiron school Enjoy SPring-8 and Enjoy Japan!